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# Determination of the $^{242}\text{Pu}$ Branching Ratio via Alpha-Gamma Coincidence

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## Abstract

When the burn-up is high, the  $^{242}\text{Pu}$  isotopic content becomes more important. The traditional correlation method <sup>1)</sup> will fail. The  $^{242}\text{Pu}$  isotopic content in the sample plays an essential role if the neutron coincidence method is used to quantify the total amount of plutonium. In one of the earlier measurements <sup>2)</sup> we had a chance to measure an isotopic pure ( $> 99.95\%$ )  $^{242}\text{Pu}$  thick sample and realized that the difference in the branching ratio (BR) value among current nuclear data <sup>3)</sup> for the two important gamma-rays at 103.5-keV and 158.8-keV. In this study, the thick sample was counted on a 15% ORTEC safeguards type HPGe to further improve BR determination of the 159-keV gamma-ray. Furthermore, we have made a thin  $^{242}\text{Pu}$  sample from the thick sample and performed alpha-gamma coincidence measurements. Our preliminary gamma-ray BR results are  $4.37(6) \times 10^{-4}$ ,  $2.79(8) \times 10^{-5}$ , and  $2.25(8) \times 10^{-6}$  for 44.9-keV, 103.5-keV, and 158.9-keV, respectively.

## Introduction

In the previous paper<sup>2</sup>, we have mentioned that we were surprised to learn the difference in the branching ratio (BR) value among current nuclear data for the two important gamma-rays at 103.5-keV and 158.8-keV. For example, the Table of Radioactive Isotopes<sup>3</sup> values are  $7.8(8) \times 10^{-5}$  and  $4.5(15) \times 10^{-6}$ , respectively; the IAEA evaluation<sup>4</sup> values are  $2.63(9) \times 10^{-5}$  and  $3.0(2) \times 10^{-6}$ , respectively. The IAEA evaluation results are predominately based on a study<sup>5</sup> by R. Vaninbroux et al. The LLNL MGA code<sup>1</sup> uses the IAEA evaluated value for 103.5-keV and since the MGA code does not use 158.8-keV for  $^{242}\text{Pu}$  determination, no information of the branching ratio of this peak is found in the code. A more recent publication<sup>4</sup> by Berlizov et. al. also reported that the BR of the 158.8-keV gamma-ray is lower by 35% at  $2.20(8) \times 10^{-6}$  via both gamma-ray measurement using a planar HPGe detector and a neutron coincidence measurement. Figure 1 shows a compiled<sup>8</sup> decay scheme of  $^{242}\text{Pu}$ . The alpha decay of  $^{242}\text{Pu}$  populates states in  $^{238}\text{U}$  at excitation energy of 0-keV, 44.91-keV, 148.41-keV, and 307.21-keV. The transition from 307-keV to 148-keV state produces gamma-ray with energy of 159-keV, etc. Because of small alpha branch to the higher excitation state and long half-life of  $^{242}\text{Pu}$  (compared to the rest of plutonium isotopes) as well as cascade summing of the gamma-rays thus made the gamma-ray branching intensity difficult to determine. Significant differences of the three gamma-ray emission probability have been reported.

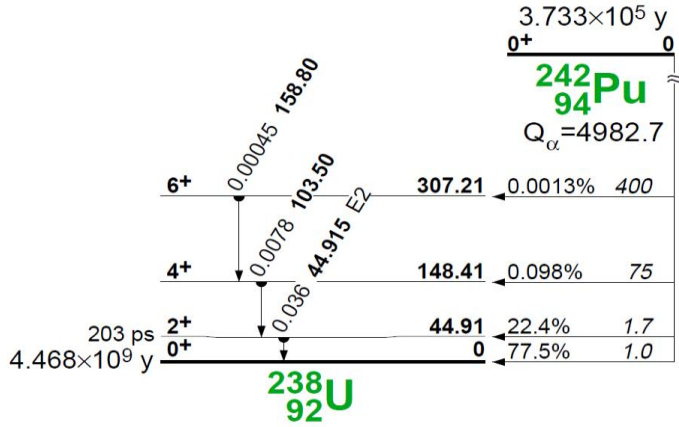


Figure 1. Alpha- and gamma-ray decay branch of  $^{242}\text{Pu}$  taken from reference 8 compilation, most of the alpha decays to the first excited state (22.4%) and ground state (77.5%) in  $^{238}\text{U}$ . Gamma-ray emission probabilities shown are per 100 decay of  $^{242}\text{Pu}$ . The gamma-ray BR values are quite different from those in Ref 4.

### Experimental setup

The  $^{242}\text{Pu}$  sample was acquired from Oakridge National Laboratory with mass spectrometry results shown in Table 1 as well as decay corrected weight percentages to the present date, the t-half used in the calculations.

	Declared (08/01/07)	Present day (02/01/12)	t-half(yr)
Total PuO <sub>2</sub> (4.53g)	Weight percentage		
238Pu	0.0032	0.0031	87.7
239Pu	0.0049	0.0049	24100
240Pu	0.0218	0.0218	6537.3
241Pu	0.0092	0.0074	14.35
242Pu	99.9589	99.9608	387000

This plutonium oxide sample (about 4.5 g in weight) was placed on the top of a ORTEC safeguards type HPGe detector (SGD-GEM-5030P4) that has energy resolution of  $\sim 650\text{eV}$  @ 120-keV and  $\sim 725\text{ eV}$  @ 160-keV. Digital data acquisition system, ORTEC DSPECpro, was used. A 30mil thick Cd absorber was placed between the sample and the HPGe detector to attenuate the 59-keV from the decay of  $^{241}\text{Am}$ . Data were collected for 10-days. To minimize

effects due to summing and dead time (about 4% during the measurement), the sample was placed on the axis 15cm from the face of the detector.

A thin  $^{242}\text{Pu}$  sample was also made by the same material evaporated on a thin platinum circular foil of  $300\text{mm}^2$  in area, 542 mg in weight, for particle-gamma measurement. The foil serves as the window to the vacuum system where an ORTEC silicon surface barrier (Si(SB)) detector of  $150\text{ }\mu\text{m}$  thick  $300\text{mm}^2$  in area was mounted with its front facing the window. The evaporated side was facing the Si(SB) detector and the non-evaporated side was facing an ORTEC planar HPGe detector (GLP-36XXX) with energy resolution of 600 eV @ 120-keV. NIM electronics were used. Amplifiers, (ORTEC 572) with shaping time of 0.5  $\mu\text{s}$  and 2  $\mu\text{s}$ , were employed for Si(SB) and HPGe detectors, respectively. List-mode data acquisition system, FAST COMTEC MPA-3, was used for event-by-event recording with a time accuracy of 50 ns. List mode data were collected at for 15-days. The time and energy information was used to extract alpha-gamma coincidence information. This coincidence measurement was used to minimize  $^{241}\text{Pu}$  (mostly beta-decay) gamma-ray at 44.20-keV, 44.86-keV and 103.68-keV contribution to the gamma-ray energy of interest at 44.92-keV and 103.50-keV. The alpha coincidence requirements also cleaned up the beta-related continuum in the planar HPGe detector. Two gamma-ray calibration standards,  $^{109}\text{Cd}$  and  $^{57}\text{Co}$ , were used to interpolate the detection efficiency at 103.5 keV. Figure 2 shows a spectrum of single gamma-ray measurement with the thick  $^{242}\text{Pu}$  sample and Figure 3 showing a 2-D coincidence matrix with gamma-ray energy and particle energy measured with the thin  $^{242}\text{Pu}$  sample.

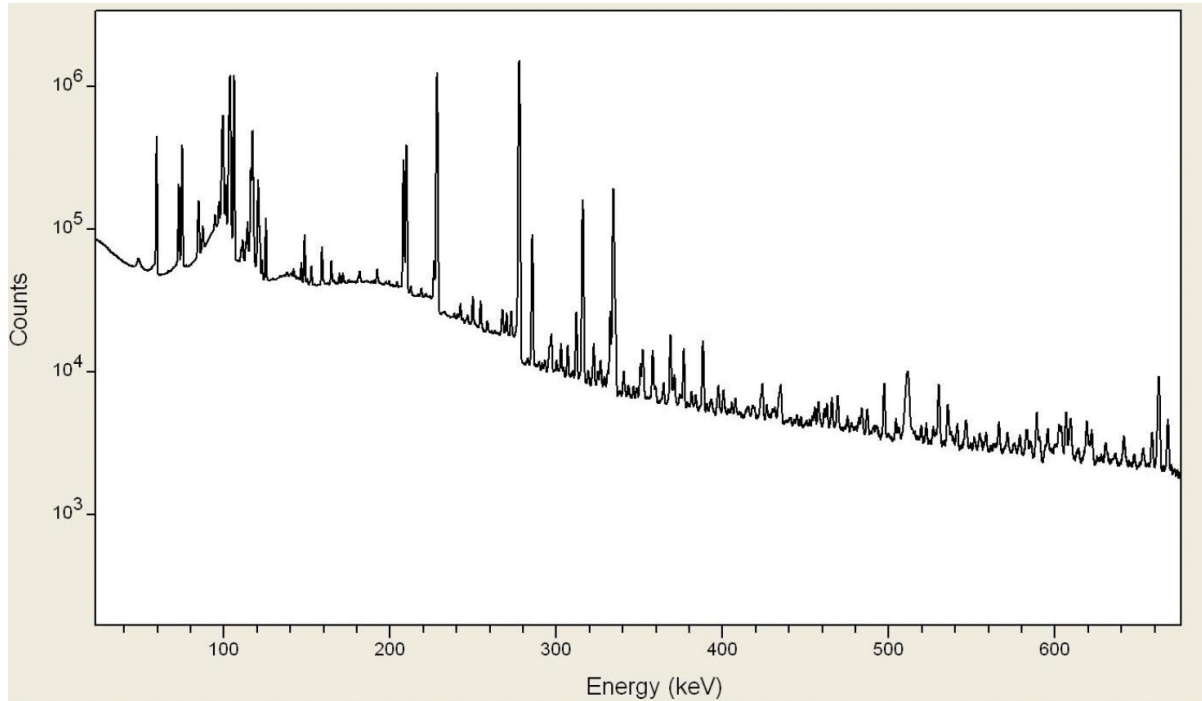


Figure 2. Gamma-ray spectrum of the thick  $^{242}\text{Pu}$  sample, 30mil Cd absorber was used to attenuate 59-keV gamma-ray from the decay of  $^{241}\text{Am}$ .

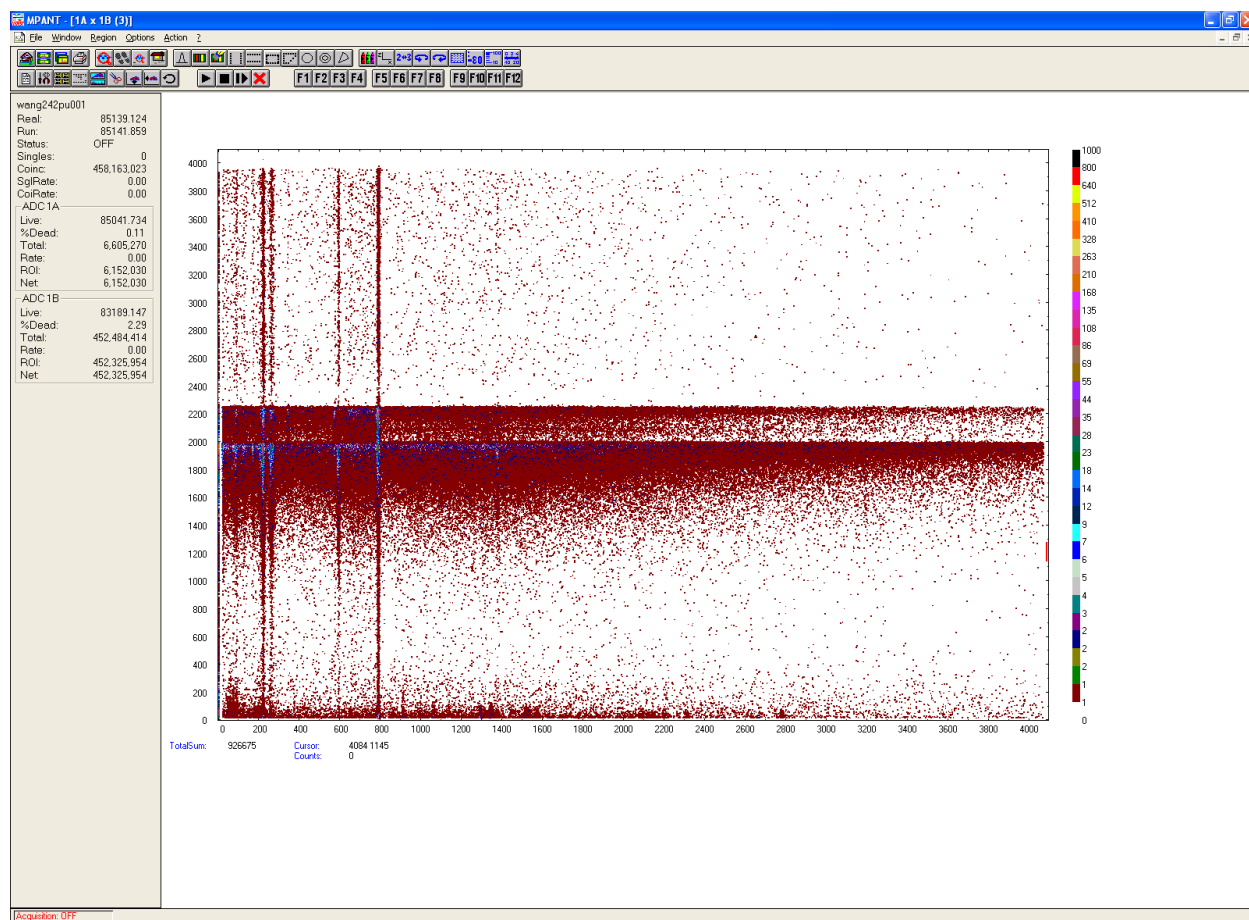


Figure 3. A 4096x4096 2D display of the particle gamma coincidence data, x-axis is gamma-ray energy and y axis is charged particle energy.

## Data Analysis and results

### a) The 158.9-keV – Thick sample, 15% efficiency safeguards detector, via relative ratios

A fit at the 158.8-keV region of the gamma-ray spectrum from the thick sample is shown in figure 4; the  $^{242}\text{Pu}$  gamma-ray branching ratios were derived using the nearby  $^{240}\text{Pu}$  and  $^{241}\text{Pu}$  gamma-ray intensities with branching ratios from an LLNL evaluation and an IAEA evaluation. The results are shown in Table 2. The energy calibration of the spectrum was performed using a mixed gamma-ray standard and measured at 158.86 (3) keV. It is important to note that this BR is sensitive to the detector resolution, the purity of  $^{242}\text{Pu}$ , the count rate and the tailing of the gamma-rays. Our preliminary result, 2.25(8) E-6, is in consistent with result from a recent measurement<sup>6)</sup> of 2.20(8) E-6.

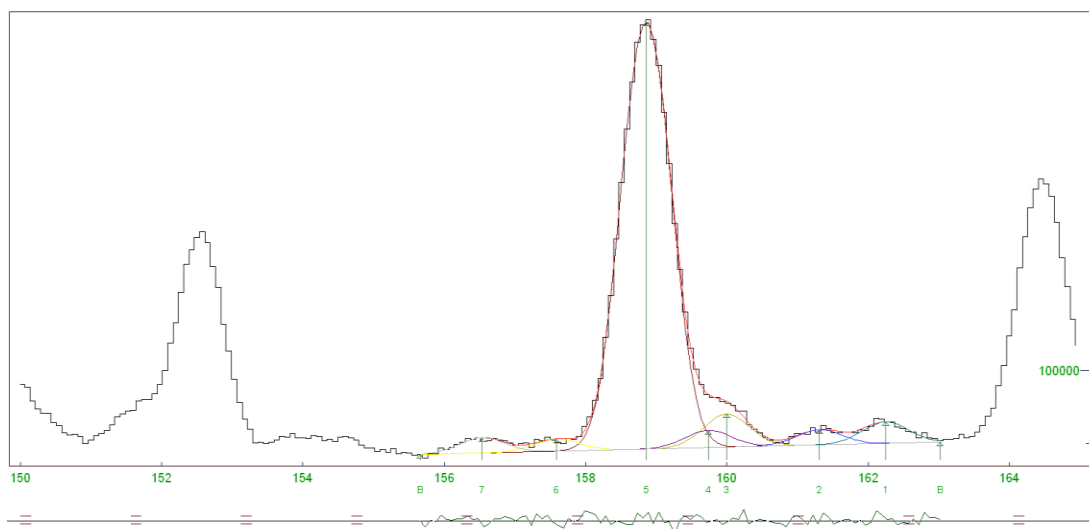


Figure 4. A fit to the 159-keV region from the thick  $^{242}\text{Pu}$  sample measurement, the residuals are plotted at the bottom

	$^{241}\text{Pu}$ BR	$^{240}\text{Pu}$ BR	$^{242}\text{Pu}(159)$ BR (derived)
LLNL(ref. 7)	6.74E-8	4.02E-6	2.25(8)E-6
IAEA (ref. 4)	6.58E-8	4.02E-6	2.29(8)E-6

Table 2. The  $^{242}\text{Pu}$  159-keV branching intensity results from this study

**b) The 44.9-keV – Thin sample, planar detector, alpha-gamma coincidence ratio determination with  $^{238}\text{Pu}$  at 43.48-keV**

Because of close geometry of the counting and the beta-decay background, alpha particle coincidence data was used with an assumption of uniform alpha coincidence efficiency across the energy of interests. The alpha coincidence efficiency was determined to be 26.7(3) % by comparing the single and coincidence data. The BR ratios of the  $^{238}\text{Pu}$  at 43.48-keV are quite similar between the IAEA evaluation and LLNL evaluation. The BR results of  $^{242}\text{Pu}(44.9\text{-keV})$  is shown in Table 3 and a fit to the coincidence data gated on alpha particles is shown in Figure 5. The energy of the  $^{238}\text{Pu}$  and  $^{242}\text{Pu}$  gamma-rays is 43.47(2) keV and 44.90(1) keV from the fit, respectively.

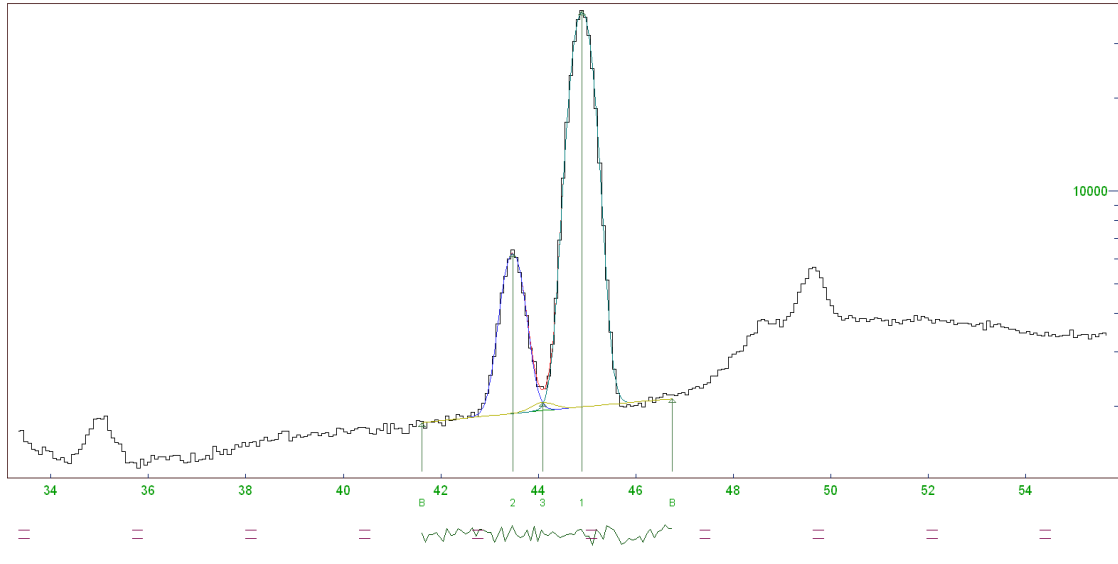


Figure 5. The 45-keV region fit of the alpha-gamma coincidence gamma-ray data gated on the alpha particles

	$^{238}\text{Pu}$ BR	$^{242}\text{Pu}(44.9)$ BR (derived)
LLNL(ref. 7)	3.92E-4	4.37(6)E-4
IAEA (ref. 4)	3.94E-4	4.39(6)E-4

Table 3 BR of 44.9-keV gamma-ray derived from two different evaluations

**c) The 103.5-keV – Thin sample, planar detector, alpha-gamma coincidence, intra- , extra- polated efficiency determination via ratio determination with 46.5-keV**

Three gamma-ray calibration standards:  $^{57}\text{Co}$ ,  $^{109}\text{Cd}$ ,  $^{241}\text{Am}$ , were used to interpolate the detector efficiency at 103.5-keV, and extrapolated to the 44.92-keV. The platinum foil is placed between these calibration standards and the detector. Using the branching ratio result from **b)** (LLNL) for 44.9-keV gamma-ray folding with proper efficiency corrections (1.75) and the fitting results of the alpha-gated gamma-ray areas, we have derived a BR of 2.79(8)E-5 for the 103.5-keV gamma-ray.

## Conclusions

Using relative BR determination via known nearby gamma-ray decay from impurity isotopes (e.g.  $^{241}\text{Pu}$  and  $^{238}\text{Pu}$ , in this study) is strongly dependent on our knowledge of the BR these impurity isotopes. The results will vary depending on the reference BR used. The particle gamma-ray coincidence technique with list-mode data acquisition significantly improves the signal to background ratio for extracting area of the gamma-rays of interest. The gating on the alpha energy also reduces interferences of the beta-decay electrons and, especially, on our sample, the  $^{243}\text{Am}$  interferences. We will improve the measurement by performing a chemical



separation of the  $^{242}\text{Pu}$  sample in the future. Table 4 shows a compilation of BR results from various studies.

Energy(keV)	Table of Isotope <sup>3)</sup>	IAEA <sup>4)</sup>	Berlizov <sup>6)</sup>	This study
44.9	3.6(5)E-4	3.72(7)E-4		4.37(6)E-4
103.5	7.8(8)E-5	2.63(9)E-5		2.79(8)E-5
158.9	4.5(15)E-6	3.0(2)E-6	2.20(8)E-6	2.25(8)E-6

Table 4 BR results from different studies.

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